The Astrophysical Journal, 705:1095–1097, 2009 November 1 doi:10.1088/0004-637X/705/1/1095

ERRATUM: "Ne ISOTOPES IN INDIVIDUAL PRESOLAR GRAPHITE GRAINS FROM THE MURCHISON METEORITE TOGETHER WITH He, C, O, Mg–Al ISOTOPIC ANALYSES AS TRACERS OF THEIR ORIGINS" (2009, ApJ, 701, 1415)

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The detection limits during Session 3 for ^{20}Ne and ^{21}Ne were interchanged. The error occurred during preparation of the manuscript, not during the laboratory experiment. The upper limits of ^{20}Ne/^{22}Ne ratios from Session 3 are based on the ^{20}Ne detection limits. The corrected Tables 2 and 3 are given below. No major conclusion in this work is affected by the revised upper limits for ^{20}Ne/^{22}Ne.

We thank an anonymous referee and Ulrich Ott for critical, constructive comments that greatly helped to improve the manuscript.

Table 2
Isotope Ratios of KFB1g Graphite Grains for Which Noble Gases have been Measured and their Most Likely Stellar Sources

Grain KFB1g-	Gas-rich?	ø (μm)	^{12}C/^{13}C	^{20}Ne/^{22}Ne	δ^{25}Mg (‰)	(^{26}Al/^{27}Al)$_{\rm inferred}$ (10^{-3})	^{16}O/^{18}O	Putative Stellar Source
64-1[b]	^{22}Ne	5.18	485.7 ± 4.4	⩽ 0.08			503.2 ± 21.2	Supernova
34-4[a]	^{20,22}Ne, ^{4}He	2.17	98.2 ± 0.8	0.032 ± 0.018	30 ± 22	6.34 ± 0.89	488.8 ± 20.8	AGB star, or less likely supernova
45-1[b]	^{20,22}Ne	3.73	15.9 ± 0.1	1.7 ± 1.0	−75 ± 33	⩽0.56	495.2 ± 18.8	Supernova?
26-2[a]		3.65	28.6 ± 0.2		24 ± 29	2.64 ± 1.25	301.8 ± 9.3	Supernova?
15-2[a]	^{22}Ne	6.18	10.4 ± 0.1	⩽ 8.6	−15 ± 22	8.17 ± 0.86	512.6 ± 18.6	Born-again AGB star? Supernova?
10-1[a]		1.94	13.7 ± 0.1		−98 ± 29	⩽0.076	508.4 ± 20.5	J-type C-star?
43-2[b]	^{22}Ne	2.36	9.0 ± 0.1	⩽ 1.1	35 ± 123	⩽2.0	529.4 ± 25.3	Born-again AGB star or supernova?
73-3[b]	^{22}Ne	1.88	8.6 ± 0.1	⩽ 6.7	39 ± 125	⩽3.7	540.7 ± 33.8	Born-again AGB star, CO nova, or Supernova?
53-2[b]		3.31	4.7 ± 0.0		101 ± 107	⩽0.98	510.5 ± 26.5	J-type C-star or CO nova?
56-1[b]		3.62	4.0 ± 0.0				491.0 ± 17.6	J-type C-star or CO nova?
57-3[b]		2.21	5.0 ± 0.0				527.7 ± 17.9	J-type C-star or CO nova?
26-3[a]	^{22}Ne	2.21	1305.1 ± 14.7	⩽ 8.6	16 ± 27	3.22 ± 0.89	512.4 ± 21.0	AGB star or supernova?
53-4[b]	^{22}Ne	2.24	246.1 ± 2.3	⩽ 8.4	−9 ± 23	1.38 ± 0.53	510.8 ± 33.8	
64-6[b]	^{22}Ne	3.18	623.3 ± 5.8	⩽ 4.8			526.8 ± 21.2	
27-5[a]	^{22}Ne	1.78	693.4 ± 8.0	⩽ 7.0	−104 ± 27	542.5 ± 33.6		
37-1[a]	^{22}Ne	3.82	857.1 ± 8.6	⩽ 8.6	6 ± 47	⩽2.8	510.3 ± 27.9	
07-1[a]		4.37	341.9 ± 3.0				529.1 ± 21.2	
12/22-1/1[a]		1.75	201.3 ± 1.7		−26 ± 22	⩽0.23	525.2 ± 19.6	
24-2[a]		1.97	512.3 ± 4.9		84 ± 53	3.92 ± 1.49	531.9 ± 25.3	
25-3[a]		2.58	965.3 ± 9.7		33 ± 22	0.16 ± 0.02	499.2 ± 24.4	
27-1[a]		2.75	87.0 ± 0.7				486.2 ± 18.3	
30-1[a]		2.58	842.6 ± 8.7		−16 ± 24	5.55 ± 1.48	553.7 ± 35.2	
30-2[a]		1.96	502.0 ± 4.7				479.3 ± 20.9	
31-1[a]		3.16	118.7 ± 1.0				513.7 ± 21.0	
31-2[a]		2.24	1064.1 ± 11.0				527.1 ± 33.4	
32-1[a]		3.05	212.1 ± 1.8				506.1 ± 19.2	
32-2/3[a]		1.89	98.3 ± 0.8		46 ± 25	1.27 ± 0.34	478.4 ± 26.2	
40-1[b]		2.65	1028.8 ± 10.4				525.2 ± 29.3	
41-1[b]		2.30	560.1 ± 5.1				514.8 ± 29.3	
42-1[b]		2.03	642.9 ± 6.2				500.2 ± 19.2	
43-1[b]		2.41	374.9 ± 3.5				516.1 ± 24.4	
43-3[b]		2.00						
45-2[b]		2.99	1159.6 ± 11.7				530.5 ± 24.9	
46-1[b]		3.77	13.0 ± 0.1			481.4 ± 17.1		
51-1[b]		2.53	147.3 ± 1.2		−7 ± 58	⩽0.43	504.8 ± 19.9	
51-2[b]		2.20	86.3 ± 0.7		43 ± 27	⩽0.15	522.4 ± 21.4	
53-1[b]		2.28	396.5 ± 3.6				515.8 ± 20.9	
57-1[b]		2.74	45.3 ± 0.4				505.0 ± 16.2	
58-1[b]		4.65	16.2 ± 0.1				447.8 ± 16.4	
61-1[b]		3.90	217.5 ± 1.9				526.3 ± 22.3	

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Table 2
(Continued)

Grain KFB1g-	Gas-rich?	ø (μm)	$^{12}C/^{13}C$	$^{20}Ne/^{22}Ne$	$\delta^{25}Mg$ (‰)	$(^{26}Al/^{27}Al)_{inferred}$ (10^{-3})	$^{16}O/^{18}O$	Putative Stellar Source
62-1[b]		4.17						
63-1[b]		3.09	87.3 ± 0.7		57 ± 25	0.62 ± 0.17	503.0 ± 18.0	
63-4[b]		2.70						
64-4[b]		1.68	453.1 ± 4.5				486.9 ± 17.3	
65-1[b]		1.84	60.8 ± 0.5				496.4 ± 24.0	
71-1[b]		4.08	592.0 ± 6.1				529.9 ± 31.0	
71-2[b]		1.87	552.5 ± 5.4				507.4 ± 22.3	
73-1[b]		3.29	1074.7 ± 12.0				524.0 ± 31.5	
73-2[b]		2.69	761.1 ± 7.9				535.7 ± 26.5	

Notes. We first list grains where one likely parent star type has been assigned, followed by grains where different types of stellar sources are possible. Uncertainties are 1σ analytical errors.
[a] Noble gas analysis session 2.
[b] Noble gas analysis session 3.

Table 3
Gas Amount and Concentrations of ^{22}Ne-rich KFB1g Graphite Grains and Upper Limits of $^{20}Ne/^{22}Ne$ Ratios of the Other Grains

Grain KFB1g-	Diameter (μm)	^{22}Ne (10^{-14} cm^3 STP)	[^{22}Ne] (10^{-3} cm^3 STP g^{-1})	$^{20}Ne/^{22}Ne$
07-1[a]	4.37			
10-1[a]	1.94			
12/22-1/1[a]	1.75			
15-2[a]	6.17	0.73 ± 0.28	0.028 ± 0.011	⩽8.6
24-2[a]	1.97			
25-3[a]	2.58			
26-2[a]	3.65			
26-3[a]	2.21	0.56 ± 0.23	0.47 ± 0.20	⩽8.6
27-1[a]	2.75			
27-5[a]	1.78	0.87 ± 0.21	1.39 ± 0.34	⩽7.0
30-1[a]	2.58			
30-2[a]	1.96			
31-1[a]	3.16			
31-2[a]	2.24			
32-1[a]	3.05			
32-2/3[a]	1.89			
34-4[a]	2.17	145.00 ± 0.71	128.0 ± 1.6	0.032 ± 0.018
37-1[a]	3.82	0.39 ± 0.23	0.064 ± 0.037	⩽8.6
40-1[b]	2.65			
41-1[b]	2.30			
42-1[b]	2.03			
43-1[b]	2.41			
43-2[b]	2.36	3.42 ± 0.12	2.340 ± 0.088	⩽1.1
43-3[b]	2.00			
45-1[b]	3.73	2.42 ± 0.12	0.419 ± 0.022	1.7 ± 1.0
45-2[b]	2.99			
46-1[b]	3.76			
51-1[b]	2.53			
51-2[b]	2.20			
53-1[b]	2.28			
53-2[b]	3.31			
53-4[b]	2.24	0.44 ± 0.10	0.354 ± 0.084	⩽8.4
56-1[b]	3.62			
57-1[b]	2.74			
57-3[b]	2.21			
58-1[b]	4.65			
61-1[b]	3.90			
62-1[b]	4.17			
63-1[b]	3.09			
63-4[b]	2.70			
64-1[b]	5.18	48.40 ± 0.25	3.140 ± 0.040	⩽0.08
64-4[b]	1.68			
64-6[b]	3.18	0.77 ± 0.11	0.215 ± 0.032	⩽4.8

Table 3
(Continued)

Grain KFB1g-	Diameter (μm)	^{22}Ne (10^{-14} cm^3 STP)	[^{22}Ne] (10^{-3} cm^3 STP g^{-1})	^{20}Ne/^{22}Ne
65-1[b]	1.84			
71-1[b]	4.08			
71-2[b]	1.87			
73-1[b]	3.29			
73-2[b]	2.69			
73-3[b]	1.88	0.55 ± 0.14	0.74 ± 0.18	$\leqslant$6.7

Notes. Uncertainties are 1σ analytical errors.
[a] Noble gas analysis session 2.
[b] Noble gas analysis session 3.